

Some Orbital Mechanics Exercises on the Acceleration of an Orbital Body

1. Consider the radial-transverse component form (see the Appendix) of the expression for the velocity \mathbf{v} of the orbital body:

$$\mathbf{v} = \frac{v_0}{1+e} [e \sin \theta \mathbf{e}_r + (1+e \cos \theta) \mathbf{e}_\theta]$$

Based upon the definition for the acceleration \mathbf{a} of the orbital body, utilize this expression for \mathbf{v} to obtain the following expression for \mathbf{a} :

$$\mathbf{a} = -\frac{\xi}{r^2} \mathbf{e}_r$$

(Hint: $\dot{\theta} = h_0/r^2$, $v_0 = \sqrt{\xi(1+e)}/r_0$, and $h_0 = \sqrt{\xi(1+e)r_0}$.)

2. Explain why this expression for \mathbf{a} is completely expected in the case of a single attractive central force with an inverse-square-law dependence.

(Hint: Consider *Newton's second law of motion*.)

3. Show that the expression for \mathbf{a} indicated above can be recast as

$$\mathbf{a} = -\frac{v_0^2}{r_0} \frac{(1+e \cos \theta)^2}{(1+e)^3} \mathbf{e}_r$$

This latter expression can be utilized to determine \mathbf{a} at any location along the orbital path (which is identified by θ) or at any instant during the orbital motion (from which θ can be evaluated by means of the final results for the solution to the *Kepler problem*).

(Hint: $r = r_0(1+e)/(1+e \cos \theta)$.)

Spoiler Warning: Exercise Solutions/Answers appear on back of sheet.

Exercise Solutions/Answers

1. Since the acceleration \mathbf{a} is (by definition) the time-derivative of the velocity \mathbf{v} :

$$\begin{aligned}
 \mathbf{a} &= \frac{d}{dt} \left\{ \frac{v_0}{1+e} [e \sin \theta \mathbf{e}_r + (1+e \cos \theta) \mathbf{e}_\theta] \right\} \\
 &= \frac{d}{d\theta} \left\{ \frac{v_0}{1+e} [e \sin \theta \mathbf{e}_r + (1+e \cos \theta) \mathbf{e}_\theta] \right\} \frac{d\theta}{dt} \\
 &= \frac{v_0}{1+e} \dot{\theta} \left[e \cos \theta \mathbf{e}_r + e \sin \theta \frac{d\mathbf{e}_r}{d\theta} - e \sin \theta \mathbf{e}_\theta + (1+e \cos \theta) \frac{d\mathbf{e}_\theta}{d\theta} \right] \\
 &= \frac{v_0}{1+e} \dot{\theta} [e \cos \theta \mathbf{e}_r + e \sin \theta \mathbf{e}_\theta - e \sin \theta \mathbf{e}_\theta - (1+e \cos \theta) \mathbf{e}_r] \\
 &= -\frac{v_0}{1+e} \dot{\theta} \mathbf{e}_r
 \end{aligned}$$

But from the Appendix, $\dot{\theta} = h_0/r^2$, $v_0 = \sqrt{\xi(1+e)/r_0}$, and $h_0 = \sqrt{\xi(1+e)r_0}$. Thus,

$$\begin{aligned}
 \mathbf{a} &= -\frac{v_0}{1+e} \frac{h_0}{r^2} \mathbf{e}_r \\
 &= -\frac{\xi(1+e)}{(1+e)r^2} \mathbf{e}_r \\
 &= -\frac{\xi}{r^2} \mathbf{e}_r
 \end{aligned}$$

2. From elementary physics, $\mathbf{F} = m \mathbf{a}$, where \mathbf{F} denotes the *total external force* that acts upon the orbital body. As a result, $\mathbf{a} = \mathbf{F}/m$. But in the case of a single attractive central force with an inverse-square-law dependence, $\mathbf{F} = -(m\xi/r^2) \mathbf{e}_r$, and so it must follow that $\mathbf{a} = -(\xi/r^2) \mathbf{e}_r$.

3. Since $r = r_0(1+e)/(1+e \cos \theta)$, and $v_0 = \sqrt{\xi(1+e)/r_0} \Rightarrow \xi = (r_0 v_0^2)/(1+e)$,

$$\begin{aligned}
 \mathbf{a} &= -\frac{r_0 v_0^2}{1+e} \frac{(1+e \cos \theta)^2}{r_0^2 (1+e)^2} \mathbf{e}_r \\
 &= -\frac{v_0^2}{r_0} \frac{(1+e \cos \theta)^2}{(1+e)^3} \mathbf{e}_r
 \end{aligned}$$

which expresses \mathbf{a} entirely as a function of θ . The factor v_0^2/r_0 is a *characteristic acceleration* for the orbital body, which has the form of a *centripetal acceleration*.